Physics 221B Academic Year 2021–22 Notes 52 Electromagnetic Interactions With the Dirac Field[†]

1. Introduction

In the previous set of notes we second-quantized the Dirac equation for a free particle, and using ideas from hole theory we created a field theory of free electrons and positrons. Notable elements of this procedure were the use of anticommuting creation and annihilation operators and the redefinition of the vacuum to be the state corresponding to the filled Dirac sea in first quantized hole theory. In this set of notes we solve some problems involving the interaction of the second quantized Dirac field with an electromagnetic field. In the first application we adopt a model in which the electromagnetic field is just a given *c*-number field interacting with the quantized Dirac field; this allows us to find the cross section for a relativistic version of Rutherford scattering, called *Mott scattering*. In a second application our model takes both the electromagnetic field and the Dirac field to be quantized, giving us a theory of electrons, positrons and photons. We use this to find the cross section for $e^+ \cdot e^-$ annihilation. In the process of working through these examples we develop several computational techniques that are useful for similar problems in quantum electrodynamics. In these notes we continue with natural units, in which $\hbar = c = 1$.

2. The Dirac Field Interacting With an External Electromagnetic Field

The Dirac Lagrangian density for the free field was given by Eq. (51.23), which we reproduce here with a slight change of notation,

$$\mathcal{L}_0 = \bar{\psi}(i\gamma^\mu \partial_\mu - m)\psi = \bar{\psi}(i\partial \!\!\!/ - m)\psi, \tag{1}$$

where the 0-subscript means "free particle." Note that $i \not = i \gamma^{\mu} \partial_{\mu}$ can also be written \not , if p_{μ} is interpreted as the 4-momentum operator $i \partial_{\mu}$ of the first quantized Dirac theory. To incorporate interactions with the electromagnetic field, we introduce the vector potential $A_{\mu}(x)$, assumed to be a given function of $x = (\mathbf{x}, t)$. We assume that the interaction between the Dirac field and the electromagnetic field is given by the minimal coupling prescription, which in the present context means making the replacement

$$i\partial_{\mu} \to i\partial_{\mu} - qA_{\mu}.$$
 (2)

[†] Links to the other sets of notes can be found at:

http://bohr.physics.berkeley.edu/classes/221/2122/221.html.

It is a guess that this gives us the correct interaction between the Dirac field and the electromagnetic field, but as we have seen minimal coupling works well in the first quantized Dirac theory of the electron. Minimal coupling causes the Lagrangian density \mathcal{L}_0 to be replaced by

$$\mathcal{L} = \mathcal{L}_0 + \mathcal{L}_1,\tag{3}$$

where the interaction Lagrangian is given by

$$\mathcal{L}_1 = -q\,\bar{\psi}\gamma^\mu A_\mu\psi = -q\,\bar{\psi}\mathcal{A}\psi = -J^\mu A_\mu,\tag{4}$$

where $J^{\mu} = q \bar{\psi} \gamma^{\mu} \psi$ is the charge current in the Dirac theory.

Altogether, the Lagrangian density is

$$\mathcal{L} = \bar{\psi}(i\partial \!\!\!/ - qA - m)\psi. \tag{5}$$

Regarded as a classical Lagrangian density for the "classical" fields ψ and $\bar{\psi}$, this gives the Euler-Lagrange equation,

$$(i\partial - qA - m)\psi = 0, (6)$$

which is the first quantized Dirac equation for a particle interacting with an external electromagnetic field. See Eq. (48.13) and the derivation of Eq. (51.25). From the Lagrangian density we compute the Hamiltonian density as in Sec. 51.10, finding

$$\mathcal{H} = \mathcal{H}_0 + \mathcal{H}_1,\tag{7}$$

where \mathcal{H}_0 , the free particle Hamiltonian density, is

$$\mathcal{H}_0 = \bar{\psi}(-i\boldsymbol{\gamma}\cdot\nabla + m)\psi = \psi^{\dagger}(-i\boldsymbol{\alpha}\cdot\nabla + m\beta)\psi, \tag{8}$$

as in Eq. (51.36), and where the interaction Hamiltonian density is

$$\mathcal{H}_1 = -\mathcal{L}_1 = q \,\bar{\psi} \gamma^\mu A_\mu \psi. \tag{9}$$

This is all at the "classical" or first quantized level.

To second-quantize the electron-positron field we interpret ψ and $\bar{\psi}$ (but not A_{μ}) as field operators and normal order. Then the Hamiltonian is

$$H = \int d^3 \mathbf{x} \left(\mathcal{H}_0 + \mathcal{H}_1 \right) = H_0 + H_1, \tag{10}$$

where

$$H_0 = \int d^3 \mathbf{x} : \psi^{\dagger} (-i\boldsymbol{\alpha} \cdot \nabla + m\beta)\psi := \sum_{ps} E_p (b_{ps}^{\dagger}b_{ps} + d_{ps}^{\dagger}d_{ps}), \tag{11}$$

exactly as in Eq. (51.85a), and where

$$H_1 = q \int d^3 \mathbf{x} : \bar{\psi} \gamma^{\mu} A_{\mu} \psi :.$$
(12)

We will apply perturbation theory to this Hamiltonian, with the free particle part H_0 serving as the unperturbed system.



Fig. 1. Feynman diagrams for the interaction of the electron-positron field with an external electromagnetic field. The external field is indicated by \times . Notice the minus sign on diagram (d), arising from the normal ordering.

Let us look at the processes engendered by H_1 in first order perturbation theory. The relevant matrix element is $\langle b|H_1|a\rangle$, where $|a\rangle$ and $|b\rangle$ are two states. The same matrix element appears in higher order perturbation theory. It has the structure

$$\langle b|H_1|a\rangle \sim \int d^3 \mathbf{x} \,\langle n| : \sum_{ps} (b_{ps}^{\dagger} \dots d_{ps}) \,\gamma^{\mu} A_{\mu}(\mathbf{x}, t) \sum_{p's'} (b_{p's'} \dots d_{p's'}^{\dagger}) : |i\rangle, \tag{13}$$

where we use Eqs. (51.81) and (51.82) and where we omit all normalization constants, etc., and concentrate just on the creation and annihilation operators. There are four combinations of creation and annihilation operators that occur in H_1 , corresponding to the four types of Feynman diagrams illustrated in Fig. 1. As in Notes 43, the Feynman diagrams are basically space-time diagrams, with time increasing from bottom to top.

At the top of the figure is the product of creation and annihilation operators corresponding to the diagram below it. Note the minus sign in part (d) of the figure, which is due to the normal ordering. Part (a) of the figure is the scattering of an electron; part (b) is electron-positron pair creation; part (c) is pair annihilation; and part (d) is the scattering of a positron, all taking place in the external field. The interaction with the external field is marked with the symbol \times . Note that all four diagrams are topologically identical, and just differ by their orientations in space-time. By convention the arrows on the electron lines point in the direction of increasing time, while those on positron lines point backwards in time. This follows the ideas of Stückelberg and Feynman in which a positron is, in a sense, an electron traveling backwards in time. This convention makes the arrows flow continuously in the same direction along a particle line, even when an electron and positron are created or are annihilated.

3. Scattering by an External Potential $\Phi(\mathbf{x})$

Let us now consider scattering of an electron by an external electrostatic potential, $\Phi(\mathbf{x})$. This

is a model, for example, of scattering by an atomic nucleus. Thus we take

$$A^{\mu} = (\Phi, \mathbf{0})$$
 and $A = \gamma^{\mu} A_{\mu} = \gamma^{0} \Phi(\mathbf{x}).$ (14)

We let the incident electron have mode (ps) and the scattered electron have mode (p's'), so the initial and final states are

$$|i\rangle = b_{ps}^{\dagger}|0\rangle, \qquad |n\rangle = b_{p's'}^{\dagger}|0\rangle. \tag{15}$$

We set q = -e for the charge of the electron. Then the matrix element appearing in first-order time-dependent perturbation theory is

$$M = \langle n|H_{1}|i\rangle = \langle 0|b_{p's'}H_{1}b_{ps}^{\dagger}|0\rangle$$

= $-\frac{e}{V}\int d^{3}\mathbf{x} \langle 0|b_{p's'}: \sum_{p_{1}s_{1}}\sqrt{\frac{m}{E_{1}}} (b_{p_{1}s_{1}}^{\dagger}\bar{u}_{p_{1}s_{1}}e^{-i\mathbf{p}_{1}\cdot\mathbf{x}} + d_{p_{1}s_{1}}\bar{v}_{p_{1}s_{1}}e^{i\mathbf{p}_{1}\cdot\mathbf{x}})$
 $\times \gamma^{0}\Phi(\mathbf{x}) \sum_{p_{2}s_{2}}\sqrt{\frac{m}{E_{2}}} (b_{p_{2}s_{2}}u_{p_{2}s_{2}}e^{i\mathbf{p}_{2}\cdot\mathbf{x}} + d_{p_{2}s_{2}}^{\dagger}v_{p_{2}s_{2}}e^{-i\mathbf{p}_{2}\cdot\mathbf{x}}): b_{ps}^{\dagger}|0\rangle,$ (16)

where we have used the Fourier series (51.81) and (51.82) for $\psi(\mathbf{x})$ and $\bar{\psi}(\mathbf{x})$, and where we have been careful to use dummy variables of summation (p_1s_1) and (p_2s_2) so as to avoid confusion with the initial and final electron modes (ps) and (p's'). Notice that Eq. (16) contains an implied multiplication of a row spinor $(\bar{u}_{p_1s_1} \text{ or } \bar{v}_{p_1s_1})$ times a Dirac matrix (γ^0) times a column spinor $(u_{p_2s_2})$ or $v_{p_2s_2}$).

Because of the anticommutation relations among the creation and annihilation operators, the only term from the two Fourier series that survives is the $b^{\dagger}b$ term with $(p_1s_1) = (p's')$ and $(p_2s_2) = (ps)$. That is, when we evaluate the field part of the matrix element (16), we get

$$\langle 0|b_{p's'}b_{p_1s_1}^{\dagger}b_{p_2s_2}b_{ps}^{\dagger}|0\rangle = \delta_{pp_2}\,\delta_{ss_2}\,\langle 0|b_{p's'}b_{p_1s_1}^{\dagger}|0\rangle = \delta_{pp_2}\,\delta_{ss_2}\,\delta_{p'p_1}\,\delta_{s's_1},\tag{17}$$

where we have used the anticommutation relations (51.53) to migrate annihilation operators b toward the vacuum ket on the right, which they annihilate. All other terms such as the one in $b^{\dagger}d^{\dagger}$ give zero, so that only the electron scattering diagram (a) in Fig. 1 contributes. Thus we find

$$M = -\frac{e}{V} \int d^3 \mathbf{x} \sqrt{\frac{m^2}{EE'}} \left(\bar{u}_{p's'} \gamma^0 u_{ps} \right) e^{-i(\mathbf{p}' - \mathbf{p}) \cdot \mathbf{x}} \Phi(\mathbf{x})$$

$$= -\frac{e}{V} (2\pi)^{3/2} \sqrt{\frac{m^2}{EE'}} \tilde{\Phi}(\mathbf{q}) \left(\bar{u}_{p's'} \gamma^0 u_{ps} \right),$$
(18)

where we use the convention (34.78) for the Fourier transform $\tilde{\Phi}$ of Φ and where $\mathbf{q} = \mathbf{p}' - \mathbf{p}$ is the momentum transfer.

According to first-order time-dependent perturbation theory, the probability P_n of the system being in final state $|n\rangle$ is

$$P_n(t) = 2\pi t \,\Delta_t(\omega_{ni}) |\langle n|H_1|i\rangle|^2,\tag{19}$$

where n, the general notation for the final state, is identified with (p's') in our application. To compute the differential cross section $d\sigma/d\Omega'$, where Ω' refers to the direction of the scattered electron, we must sum over all final states (p's') such that \mathbf{p}' lies in a small cone of solid angle $d\Omega'$ centered on some direction (θ', ϕ') . We must also divide by the time to get the transition rate, and by the incident flux, which is

$$J_{\rm inc} = \frac{v}{V} = \frac{p_3}{EV},\tag{20}$$

where $v, p_3 = |\mathbf{p}|$ and E are the velocity, momentum and energy of the incident electron. Because of box normalization, there is one incident electron in the box of volume V, so that the density of incident particles is 1/V. We write p_3 for the magnitude of the incident 3-momentum \mathbf{p} since we are using the symbol $p = (E, \mathbf{p})$ for the 4-momentum. The initial and final energies are E and E', so ω_{ni} in Eq. (20) is identified with E' - E.

We follow the usual procedures in time-dependent perturbation theory, including replacing $\Delta_t(\omega_{ni})$ by $\delta(E - E')$, and making the replacement,

$$\sum_{\mathbf{p}'\in\text{cone}} \to \frac{V}{(2\pi)^3} \int_{\text{cone}} d^3 \mathbf{p}' = \frac{V}{(2\pi)^3} \Delta \Omega' \int_0^\infty p_3'^2 dp_3',\tag{21}$$

where $p'_3 = |\mathbf{p}'|$ is the magnitude of the final 3-momentum, not to be confused with the final 4momentum $p' = (E', \mathbf{p}')$. Then

$$\frac{d\sigma}{d\Omega'} = \frac{EV}{p_3} \frac{V}{(2\pi)^3} \int_0^\infty p_3'^2 dp_3' 2\pi \,\delta(E' - E) \,|M|^2.$$
(22)

Now using the rules of δ -functions we can write

$$\delta(E' - E) = \frac{E'}{p'_3} \delta(p'_3 - p_3), \tag{23}$$

and using Eq. (18) for M we obtain finally

$$\frac{d\sigma}{d\Omega'} = 2\pi e^2 m^2 |\tilde{\Phi}(\mathbf{q})|^2 |\bar{u}_{p's'}\gamma^0 u_{ps}|^2.$$
(24)

Note that the δ -function forces $p'_3 = p_3$ and therefore E' = E, so that after the integral has been done the initial and final 4-momenta become

$$p = (E, \mathbf{p}), \qquad p' = (E, \mathbf{p}'),$$
 (25)

that is, with the same energy.

4. Summing and Averaging over Spin States

The differential cross section (22) depends the energy E of the incident momentum, the directions of the initial and final momenta \mathbf{p} and \mathbf{p}' , and the spins s and s' of the initial and final electron states. Recall that in the nonrelativistic (electrostatic) limit the scattering cross section is

independent of the spin. (See Eq. (34.80), in which \tilde{U} can be replaced by $-e\tilde{\Phi}$ to compare with Eq. (22).)

In many applications the initial electron beam is unpolarized and we do not care about the spin of the scattered electron. Then the effective cross section is obtained by averaging over initial spin states and summing over final spin states. This means that we must make the replacement,

$$|\bar{u}_{p's'}\gamma^0 u_{ps}|^2 \to \frac{1}{2} \sum_{ss'} |\bar{u}_{p's'}\gamma^0 u_{ps}|^2.$$
 (26)

Spin sums of this type are very common in problems involving relativistic electrons, so we will now illustrate some techniques that are useful in evaluating them.

First we write

$$\frac{1}{2}\sum_{ss'}|\bar{u}_{p's'}\gamma^0 u_{ps}|^2 = \frac{1}{2}\sum_{ss'}(\bar{u}_{p's'}\gamma^0 u_{ps})(\bar{u}_{ps}\gamma^0 u_{p's'}).$$
(27)

This result would be more obvious if the second factor were written

$$(\bar{u}_{p's'}\gamma^0 u_{ps})^*,\tag{28}$$

but in fact they are the same, since the expression (28) can be written

$$(u_{p's'}^{\dagger}\gamma^{0}\gamma^{0}u_{ps})^{*} = u_{ps}^{\dagger}\gamma^{0}\gamma^{0}u_{p's'} = \bar{u}_{p's'}\gamma^{0}u_{p's'},$$
(29)

since γ^0 is Hermitian. Now the sum (27) can be written as a trace,

$$\frac{1}{2}\operatorname{tr}\Big[\gamma^0\Big(\sum_s u_{ps}\bar{u}_{ps}\Big)\gamma^0\Big(\sum_{s'} u_{p's'}\bar{u}_{p's'}\Big)\Big],\tag{30}$$

in which the sums are positive energy projectors Π_+ , as indicated by Eq. (50.146). Then we can use Eq. (50.38) to write the spin sum as

In this manner, the spin sum has been transformed into the trace of a polynomial in the γ matrices.

5. Traces of Products of γ -Matrices

We now present some rules for evaluating the traces of products of γ -matrices. The rules we will use are

$$\operatorname{tr} 1 = 4, \tag{32a}$$

$$\operatorname{tr}(\gamma^{\mu}\gamma^{\nu}) = 4g^{\mu\nu},\tag{32b}$$

$$\operatorname{tr}(\gamma^{\mu}\gamma^{\nu}\gamma^{\alpha}\gamma^{\beta}) = 4(g^{\mu\nu}g^{\alpha\beta} - g^{\mu\alpha}g^{\nu\beta} + g^{\mu\beta}g^{\nu\alpha}), \qquad (32c)$$

 $tr(any odd number of \gamma's) = 0.$ (32d)

Alternative versions of rules (32b) and (32c) are

$$\operatorname{tr}(\not a \not b) = 4(a \cdot b), \tag{33a}$$

$$\operatorname{tr}(\mathfrak{a}\mathfrak{b}\mathfrak{a}\mathfrak{a}) = 4[(a \cdot b)(c \cdot d) - (a \cdot c)(b \cdot d) + (a \cdot d)(b \cdot c)], \tag{33b}$$

where we use the notation $(a \cdot b) = a_{\mu}b^{\mu}$ for the scalar product of two 4-vectors, as in Eq. (47.104).

Now we present the proofs of these formulas. The proof of Eq. (32a) is trivial. We prove Eq. (32b) with the help of the fundamental anticommutation relation (48.18),

$$\operatorname{tr}(\gamma^{\mu}\gamma^{\nu}) = 2g^{\mu\nu}\operatorname{tr}(1) - \operatorname{tr}(\gamma^{\nu}\gamma^{\mu}).$$
(34)

But by cycling the products in the final trace, it becomes equal to the trace on the left-hand side. Using Eq. (32a) we then obtain Eq. (32b).

The proof of Eq. (32c) is similar but more elaborate. We use the fundamental anticommutation relation (48.18) to migrate γ^{μ} to the right in three steps, obtaining

$$\operatorname{tr}(\gamma^{\mu}\gamma^{\nu}\gamma^{\alpha}\gamma^{\beta}) = 2g^{\mu\nu}\operatorname{tr}(\gamma^{\alpha}\gamma^{\beta}) - 2g^{\mu\alpha}\operatorname{tr}(\gamma^{\nu}\gamma^{\beta}) + 2g^{\mu\beta}\operatorname{tr}(\gamma^{\nu}\gamma^{\alpha}) - \operatorname{tr}(\gamma^{\nu}\gamma^{\alpha}\gamma^{\beta}\gamma^{\mu}), \quad (35)$$

whereupon by cycling the factors in the final trace it becomes equal to the left-hand side. Then with the aid of Eq. (32b) we derive Eq. (32c). It is clear that with techniques like this the trace of the product of any even number n of γ -matrices can be reduced sums of traces of products of n-2 γ -matrices.

The proof of Eq. (32d) uses the properties of the matrix $\gamma_5 = i\gamma^0\gamma^1\gamma^2\gamma^3$, namely, $\gamma_5^2 = 1$ and $\{\gamma_5, \gamma^{\mu}\} = 0$ (see Eqs. (48.105)). For example, in the case of the trace of the product of three γ -matrices we have

$$tr(\gamma^{\mu}\gamma^{\nu}\gamma^{\alpha}) = tr(\gamma_{5}^{2}\gamma^{\mu}\gamma^{\nu}\gamma^{\alpha}) = -tr(\gamma_{5}\gamma^{\mu}\gamma^{\nu}\gamma^{\alpha}\gamma_{5}) = -tr(\gamma_{5}^{2}\gamma^{\mu}\gamma^{\nu}\gamma^{\alpha})$$

$$= -tr(\gamma^{\mu}\gamma^{\nu}\gamma^{\alpha}) = 0,$$
(36)

where in the second equality we have migrated one of the γ_5 's past the three other γ -matrices, incurring three minus signs, and in the third we have cycled the factors, restoring the original trace but with a minus sign. This procedure obviously works for any odd number of γ -matrices.

6. The Mott Cross Section

Now we may evaluate the trace (31). Multiplying the factors out and dropping terms that are of odd order in γ^{μ} , the expression (31) becomes

$$\frac{1}{8m^2}\operatorname{tr}\left(m^2\gamma^0\gamma^0 + \gamma^0\not\!\!\!p\gamma^0\not\!\!p'\right) = \frac{1}{2m^2}\left(m^2g^{00} + 2EE' - g^{00}(p\cdot p')\right). \tag{37}$$

But $g^{00} = 1$, $p \cdot p' = EE' - \mathbf{p} \cdot \mathbf{p}'$, E = E', and $m^2 = E^2 - |\mathbf{p}|^2$, so this expression becomes

$$\frac{1}{2m^2}(2E^2 - |\mathbf{p}|^2 + \mathbf{p} \cdot \mathbf{p}') = \frac{1}{2m^2}[2E^2 - |\mathbf{p}|^2(1 - \cos\theta)],$$
(38)

where we have used $|\mathbf{p}| = |\mathbf{p}'|$ and where θ is the angle between \mathbf{p} and \mathbf{p}' , that is, it is the scattering angle. This can also be written

$$\frac{1}{m^2} [E^2 - |\mathbf{p}|^2 \sin^2(\theta/2)] = \frac{E^2}{m^2} [1 - v^2 \sin^2(\theta/2)],$$
(39)

where $v = |\mathbf{p}|/E$ is the velocity of the incident electron.

Now we can make the replacement (26) in the cross section (24), obtaining,

$$\frac{d\sigma}{d\Omega'} = 2\pi e^2 E^2 |\tilde{\Phi}(\mathbf{q})|^2 [1 - v^2 \sin^2(\theta/2)].$$
(40)

If the potential $\Phi(\mathbf{x})$ arises from an extended charge distribution, for example, the interior of a proton or a nucleus, then the method of form factors can be used, as discussed in Sec. 34.17. But in the case of scattering by a point charge Ze, for which the potential is $\Phi(\mathbf{x}) = -Ze^2/|\mathbf{x}|$, the Fourier transform of the potential is

$$\tilde{\Phi}(\mathbf{q}) = -\frac{2Ze^2}{\sqrt{2\pi}} \frac{1}{q^2}.$$
(41)

A direct evaluation of the Fourier transform of the Coulomb potential is tricky because of the singularity, but by multiplying the potential by $e^{-\kappa r}$ it becomes a Yukawa potential whose Fourier transform is easier to evaluate, and which is given in Eq. (34.121). Taking the limit $\kappa \to 0$ then gives Eq. (41). Also, the square of the 3-momentum transfer can be written,

$$q^{2} = 2|\mathbf{p}|^{2}(1 - \cos\theta) = 4|\mathbf{p}|^{2}\sin^{2}(\theta/2), \qquad (42)$$

so that altogether we obtain the cross section for the scattering of relativistic electrons by a positive point charge,

$$\frac{d\sigma}{d\Omega'} = \frac{Z^2 e^4 E^2}{4p^4 \sin^4(\theta/2)} [1 - v^2 \sin^2(\theta/2)],$$
(43)

where we now just write p for the magnitude of the 3-momentum \mathbf{p} . This is a relativistic version of the Rutherford cross section, given by Eq. (34.113); in comparing the two, it should be noted that Ein the nonrelativistic formula is $p^2/2m$, whereas in Eq. (43) it is the relativistic energy, $(m^2 + p^2)^{1/2}$, and that moreover the relativistic formula is expressed in units in which $\hbar = c = 1$. Formula (43) is called the *Mott cross section*. The correction factor $v^2 \sin^2(\theta/2)$ comes from the spin of the electron; it is absent in the scattering of charged particles of spin 0. It obviously becomes important at large scattering angles $(\theta \to \pi)$ and at high velocities $(v \to 1)$, in the formula; $v \to c$ in ordinary units). As $v \to c$, magnetic forces become as important as electric forces, which is what we see here.

7. Covariant Lagrangian for the Interacting Dirac and Electromagnetic Fields

Now we turn to problems that involve the second-quantized Dirac field and the quantized electromagnetic field, that is, problems with electrons, positrons and photons.

We begin at the "classical" level at which this system is the first-quantized Dirac field interacting with a classical or *c*-number electromagnetic field. The classical Lagrangian density is

$$\mathcal{L} = \mathcal{L}_D + \mathcal{L}_{\rm em} + \mathcal{L}_{\rm int}, \tag{44}$$

where \mathcal{L}_D is the free Dirac Lagrangian density,

$$\mathcal{L}_D = \bar{\psi}(i\partial \!\!\!/ - m)\psi \tag{45}$$

(see Eq. (51.23)), where \mathcal{L}_{em} is the Lagrangian density for the free electromagnetic field,

$$\mathcal{L}_{\rm em} = \frac{1}{16\pi} F^{\mu\nu} F_{\nu\mu} = \frac{1}{8\pi} (E^2 - B^2) \tag{46}$$

(see Eq. (51.13)), and where \mathcal{L}_{int} is the interaction Lagrangian density,

$$\mathcal{L}_{\rm int} = -J^{\mu}A_{\mu} \tag{47}$$

(see Eq. (51.19)). This is a combination of the Lagrangian density for the electromagnetic field driven by a matter current J^{μ} , as in Sec. 51.7, and the Lagrangian density (5) for the "classical" Dirac field driven by an external electromagnetic field via minimal coupling. Note that the current is given by

$$J^{\mu} = -e\bar{\psi}\gamma^{\mu}\psi, \tag{48}$$

so that

$$\mathcal{L}_{\rm int} = e\bar{\psi}\mathcal{A}\psi,\tag{49}$$

where we set q = -e for the electron.

If A^{μ} (all four components), ψ and $\bar{\psi}$ are regarded as independent fields, then the Euler-Lagrange equations arising from the Lagrangian density (44) are

$$(i\partial \!\!\!/ - m)\psi = -eA\!\!\!/\psi, \tag{50a}$$

$$\partial_{\nu}F^{\nu\mu} = 4\pi J^{\mu},\tag{50b}$$

that is, they are the first quantized Dirac equation driven by the electromagnetic field, and Maxwell's equations, driven by the Dirac charge current. Here we present Maxwell's equations in covariant form, see Eq. (E.93), because we wish to emphasize that that the equations of motion, taken as a whole, are covariant. But Eq. (50b) is equivalent to the (3 + 1)-version of Maxwell's equations, Eqs. (51.21).

8. Coulomb Gauge and Covariant Quantization

In preparation for quantization of this system we now introduce Coulomb gauge, which breaks both gauge invariance and manifest Lorentz covariance. This means that the intermediate steps in our subsequent calculations will not be Lorentz covariant, although the final answers will be, since the theory itself is Lorentz covariant and a choice of gauge is just a convention that drops out when physical answers are derived.

As pointed out in Notes 40 and shown in detail in Notes 42, one of the advantages of Coulomb gauge is that it gives good unperturbed Hamiltonians for the interaction of matter with radiation when the matter is essentially nonrelativistic and all velocities are small compared to the speed of

light. In the applications we are about to consider, however, we will be dealing with problems in which the particles may take on relativistic velocities, and ones in which particles are created and/or destroyed. Therefore this would be a good point to switch over to completely covariant methods, were it not for the extra formalism required to do so. Nevertheless, we will say a few words about what is involved in covariant quantization.

The quantization of the Dirac field outlined in Notes 51 is already fully covariant, but that of the electromagnetic field, carried out in Coulomb gauge in Notes 41, is not. The electromagnetic field is a gauge field with nonphysical gauge degrees of freedom. These are associated with "constraints," that is, algebraic relations among the field variables that are not evolution equations. The principal constraint in electromagnetism is Gauss's law, $\nabla \cdot \mathbf{E} = 4\pi\rho$, which ties the longitudinal component of the electric field to the matter degrees of freedom. It is not an evolution equation. There are fully covariant ways of quantizing the electromagnetic field, but they involve the introduction of new degrees of freedom, corresponding to "longitudinal" and "scalar" photons. Photons that correspond to actual light waves are always transverse, and these are the only ones that we see in Coulomb gauge. But in a covariant quantization, the longitudinal and scalar photons appear when describing electromagnetic fields that do not propagate, such as electrostatic fields, and they also appear in a covariant description of intermediate states in perturbation theory, that is, as virtual photons.

Moreover, the version of perturbation theory that we have been using, which is sometimes called "old fashioned perturbation theory," is not covariant. Covariant perturbation theory exists, but it is another piece of formalism we would have to master before turning to a fully covariant treatment of the interaction of the Dirac field with the electromagnetic field.

We shall proceed with Coulomb gauge and noncovariant perturbation theory. It is not wrong, and it is not too bad for the problems we shall consider. You will be able to see how the results we obtain can be assembled into covariant answers, and it how the pieces of noncovariant perturbation theory can be assembled into their fully covariant versions.

9. Coulomb Gauge in the "Classical" Theory

Recall that in Coulomb gauge the scalar potential Φ is not an independent variable, but rather it is determined by the charge density via Coulomb's law. That is,

$$\Phi(\mathbf{x}) = \int d^3 \mathbf{x}' \frac{\rho(\mathbf{x}')}{|\mathbf{x} - \mathbf{x}'|},\tag{51}$$

where in the present context the charge density is that of the Dirac electron,

$$\rho(\mathbf{x}) = -e\psi^{\dagger}(\mathbf{x})\psi(\mathbf{x}). \tag{52}$$

Recall also that the fields in Eq. (51) are evaluated at the same time, that is, Φ is the nonretarded potential. Also, the vector potential **A** is purely transverse, so its longitudinal component vanishes, and the longitudinal and transverse parts of the electric field are

$$\mathbf{E}_{\parallel} = -\nabla\Phi, \qquad \mathbf{E}_{\perp} = -\frac{\partial\mathbf{A}}{\partial t}.$$
 (53)

In Coulomb gauge it is convenient to transform the Lagrangian, the spatial integral of the Lagrangian density (44). Ignoring the Dirac part of the Lagrangian for the moment, we have

$$L_{\rm em} + L_{\rm int} = \int d^3 \mathbf{x} \left[\frac{1}{8\pi} (E^2 - B^2) - \rho \Phi + \mathbf{J} \cdot \mathbf{A} \right], \tag{54}$$

where ρ is given by Eq. (52). First we note that

$$\int d^3 \mathbf{x} \, E^2 = \int d^3 \mathbf{x} \, (\mathbf{E}_{\parallel} + \mathbf{E}_{\perp})^2 = \int d^3 \mathbf{x} \, (E_{\parallel}^2 + E_{\perp}^2), \tag{55}$$

since the cross terms involving $\mathbf{E}_{\parallel} \cdot \mathbf{E}_{\perp}$ vanish when integrated over all space (see Eq. (40.25)). As for the longitudinal electric field energy, it can be transformed as in Sec. 40.18,

$$\frac{1}{8\pi} \int d^3 \mathbf{x} \, E_{\parallel}^2 = \frac{1}{8\pi} \int d^3 \mathbf{x} \, (\nabla \Phi)^2 = \frac{1}{8\pi} \int d^3 \mathbf{x} \, [\nabla \cdot (\Phi \nabla \Phi) - \Phi \nabla^2 \Phi] = \frac{1}{2} \int d^3 \mathbf{x} \, \Phi \rho, \tag{56}$$

where we drop boundary terms and use $\nabla^2 \Phi = -4\pi\rho$ (valid in Coulomb gauge). The final expression cancels one half of the term $-\rho\Phi$ in the Lagrangian (54), so that overall the Lagrangian (including now the Dirac term) is

$$L = \int d^3 \mathbf{x} \left[\frac{1}{8\pi} (E_\perp^2 - B^2) - \frac{1}{2} \rho \Phi + \mathbf{J} \cdot \mathbf{A} + \bar{\psi} (i \partial \!\!\!/ - m) \psi \right].$$
(57)

The independent fields in this Lagrangian are the transverse components of \mathbf{A} (which has no longitudinal component), and ψ and $\overline{\psi}$. The scalar potential Φ depends on ψ and $\overline{\psi}$ through Eqs. (51)–(52), and is not independent. It can be shown that when the action $S = \int L dt$ is forced to be stationary with respect to variations in these independent fields, the result is the equations of motion (50). The derivation is slightly tricky, and is relegated to Prob. 1.

We can now derive the Hamiltonian from the Lagrangian (57). The momentum conjugate to ψ is

$$\pi = \frac{\partial \mathcal{L}}{\partial \dot{\psi}} = \bar{\psi}(i\gamma^0),\tag{58}$$

as in Eq. (51.34), while the momentum $\bar{\pi}$ conjugate to $\bar{\psi}$ vanishes. The momentum conjugate to **A** is

$$\mathbf{P} = \frac{\partial \mathcal{L}}{\partial \dot{\mathbf{A}}} = -\frac{\mathbf{E}_{\perp}}{4\pi},\tag{59}$$

where it is understood that \mathbf{E}_{\perp} is an abbreviation for $-\dot{\mathbf{A}}$ in the Lagrangian (57). Altogether, then, the Hamiltonian density is

$$\mathcal{H} = \bar{\psi} \left(i \gamma^0 \frac{\partial \psi}{\partial t} \right) + \mathbf{P} \cdot \dot{\mathbf{A}} - \mathcal{L} = \mathcal{H}_D + \mathcal{H}_{em} + \mathcal{H}_{Coul} + \mathcal{H}_T, \tag{60}$$

where

$$\mathcal{H}_D = \bar{\psi} \big(-i\boldsymbol{\gamma} \cdot \nabla + m \big) \psi = \psi^{\dagger} \big(-i\boldsymbol{\alpha} \cdot \nabla + m\beta \big) \psi, \tag{61a}$$

$$\mathcal{H}_{em} = \frac{E_\perp^2 + B^2}{8\pi},\tag{61b}$$

$$\mathcal{H}_{\text{Coul}} = \frac{1}{2}\rho\Phi,\tag{61c}$$

$$\mathcal{H}_T = -\mathbf{J} \cdot \mathbf{A} = e\bar{\psi}\boldsymbol{\gamma} \cdot \mathbf{A}\psi = e\psi^{\dagger}\boldsymbol{\alpha} \cdot \mathbf{A}\psi, \qquad (61d)$$

and where the T on \mathcal{H}_T stands for "transverse."

10. The Quantum System

We quantize the system by interpreting the fields ψ , $\bar{\psi}$ and **A** as quantum fields with the Fourier expansions (51.81), (51.82) and (41.22), respectively. All three of these employ box normalization in a box of volume V. For reference we write out the last of these in natural units ($\hbar = c = 1$):

$$\mathbf{A}(\mathbf{x}) = \sqrt{\frac{2\pi}{V}} \sum_{\lambda} \frac{1}{\sqrt{\omega}} [a_{\lambda} \hat{\boldsymbol{\epsilon}}_{\lambda} e^{i\mathbf{k}\cdot\mathbf{x}} + a_{\lambda}^{\dagger} \hat{\boldsymbol{\epsilon}}_{\lambda}^{*} e^{-i\mathbf{k}\cdot\mathbf{x}}], \tag{62}$$

where as usual $\lambda = (\mathbf{k}\mu)$ is an abbreviation for the mode of the electromagnetic field. The Hamiltonian is then the spatial integral of the Hamiltonian density, which we must normal order. We write it as $H = H_0 + H_1$, where $H_0 = H_D + H_{em}$ is the free field Hamiltonian,

$$H_D = \int d^3 \mathbf{x} : \psi^{\dagger} (-i\boldsymbol{\alpha} \cdot \nabla + m) \psi := \sum_{ps} E_p (b_{ps}^{\dagger} b_{ps} + d_{ps}^{\dagger} d_{ps}), \tag{63a}$$

$$H_{\rm em} = \int d^3 \mathbf{x} \, \frac{:E_{\perp}^2 + B^2:}{8\pi} = \sum_{\lambda} \omega_{\lambda} \, a_{\lambda}^{\dagger} a_{\lambda}, \tag{63b}$$

and where $H_1 = H_{Coul} + H_T$ is the perturbing Hamiltonian,

$$H_{\text{Coul}} = \frac{1}{2} \int d^3 \mathbf{x} : \rho(\mathbf{x}) \Phi(\mathbf{x}) := \frac{1}{2} \int d^3 \mathbf{x} \, d^3 \mathbf{x}' \, \frac{:\rho(\mathbf{x})\rho(\mathbf{x}'):}{|\mathbf{x} - \mathbf{x}'|}$$
$$= \frac{e^2}{2} \int d^3 \mathbf{x} \, d^3 \mathbf{x}' \, \frac{:[\bar{\psi}(\mathbf{x})\gamma^0\psi(\mathbf{x})][\bar{\psi}(\mathbf{x}')\gamma^0\psi(\mathbf{x}')]:}{|\mathbf{x} - \mathbf{x}'|}, \tag{64a}$$

$$H_T = -\int d^3 \mathbf{x} : \mathbf{J} \cdot \mathbf{A} := e \int d^3 \mathbf{x} : \bar{\psi}(\mathbf{x}) \boldsymbol{\gamma} \cdot \mathbf{A}(\mathbf{x}) \psi(\mathbf{x}) :.$$
(64b)

Let us now consider the processes engendered by H_T in first order perturbation theory. For simplicity we treat only H_T and ignore H_{Coul} . The matrix element has the structure

$$\langle b|H_1|a\rangle \sim \int d^3 \mathbf{x} \,\langle b| : \sum_{ps} (b_{ps}^{\dagger} \dots d_{ps}) \,\sum_{\lambda} (a_{\lambda} \dots a_{\lambda}^{\dagger}) \,\sum_{p's'} (b_{p's'} \dots d_{p's'}^{\dagger}) : |a\rangle, \tag{65}$$

as in Eq. (13) except that now the vector potential \mathbf{A} is a quantum field. There are altogether $2^3 = 8$ combinations of creation and annihilation operators, which produce the Feynman diagram fragments seen in Fig.2.

Diagrams (a1) and (d1) are the emission of a photon by an electron or positron, while (a2) and (d2) are the absorption of a photon by an electron or positron, respectively. Diagram (b1) is the creation of an electron-positron pair plus a photon out of the vacuum, and (c2) is the annihilation of the same into the vacuum. Finally, diagram (c1) is electron-positron pair. As in Fig. 1, all diagrams are topologically identical, and differ only by the directions in which the lines go in space-time.



Fig. 2. Feynman diagrams generated by H_T . These are the same as in Fig. 1, except for the addition of a photon line, representing either the creation or the destruction of a photon.

Note also that in all diagrams, charge is conserved at the vertex. As we will see, in noncovariant perturbation theory, 3-momentum is also conserved at each vertex but not energy, while in covariant perturbation theory, 4-momentum (including energy) is conserved at each vertex.

11. Electron-Positron Annihilation

Now we consider the reaction,

$$e^- + e^+ \to \gamma\gamma,$$
 (66)

that is, electron-positron annihilation into two photons, as an example of a process involving the Dirac field with multiple particles. Although there is a diagram for the annihilation of an electron and positron into a single photon, (c1) in Fig. 2, single-photon annihilation cannot take place in vacuum because it is impossible to satisfy overall energy and momentum conservation. But annihilation into two photons can occur, and we shall now calculate the cross-section for the process. We denote the modes of the incident electron and positron by (ps) and (p's') respectively, and of the two photons by $\lambda = (\mathbf{k}\mu)$ and $\lambda' = (\mathbf{k}'\mu')$. We will write 4-vectors $p = (E, \mathbf{p})$ and $p' = (E', \mathbf{p}')$ for the initial electron and positron, respectively, and 4-vectors $k = (\omega, \mathbf{k})$ and $k' = (\omega', \mathbf{k}')$ for the two final photons.

The process (66) cannot happen in first-order perturbation theory, because H_T is capable of creating only a single photon. That is,

$$\langle n|H_T|i\rangle = 0. \tag{67}$$

Therefore we will have to take H_T in second order perturbation theory to find a contribution to the process (66). As for H_{Coul} , it does not contain any photon creation operators, so it does not contribute and we will ignore it for this calculation.

The transition probability is

$$P_n = 2\pi t \,\Delta_t(\omega_{ni}) \,|M|^2,\tag{68}$$

in a general notation in which $|i\rangle$ and $|n\rangle$ are the initial and final states, and where for this problem we have

$$M = \sum_{k} \frac{\langle n | H_T | k \rangle \langle k | H_T | i \rangle}{E_i - E_k},\tag{69}$$

where $|k\rangle$ is an intermediate state. For this problem we write

$$|i\rangle = b_{ps}^{\dagger} d_{p's'}^{\dagger} |0\rangle, \qquad |n\rangle = a_{\lambda}^{\dagger} a_{\lambda'}^{\dagger} |0\rangle$$
(70)

so that

$$\omega_{ni} = \omega + \omega' - E - E'. \tag{71}$$

It is important to recognize that the final state is identified by the two modes $\{\lambda, \lambda'\}$ without regard to order, since the photons are identical. That is, the state (λ, λ') is identical to (λ', λ) .

The sum over intermediate states in principle runs over a complete set of states of both the electron-positron and the electromagnetic field, including any number of particles. But most terms in this vast sum vanish. To see which terms do contribute, let us write out the structure of the product of matrix elements that appears in Eq. (69):

$$\langle n|H_T|k\rangle\langle k|H_T|i\rangle \sim \int d^3\mathbf{x} \, d^3\mathbf{x}' \langle n|:(b^{\dagger}\dots d)(a\dots a^{\dagger})(b\dots d^{\dagger}):|k\rangle\langle k|:(b^{\dagger}\dots d)(a\dots a^{\dagger})(b\dots d^{\dagger}):|i\rangle,$$
⁽⁷²⁾

where the first matrix element contains fields $\bar{\psi} \mathbf{A} \psi$ evaluated at \mathbf{x} , and the second fields $\bar{\psi} \mathbf{A} \psi$ evaluated at \mathbf{x}' .

There are $2^6 = 64$ combinations of creation and annihilation operators in this, but most of them do not contribute. First of all, the initial state has no photons and the final state has two, so each application of the vector potential **A** must create a single photon. That is, the terms involving the a_{λ} operators do not contribute, and we keep only the ones involving a_{λ}^{\dagger} . (Recall that time advances from right to left, taking us from the initial state to the final state.) Next, because we must annihilate the incident e^- , e^+ pair in two steps, and because there are a total of four fermion creation or annihilation operators between the initial and final state, we must use three annihilation and one creation operator. Moreover, the pattern must be AAAC or AACA, because, if it were, for example, CAAA, then the intermediate state would contain no fermions, and the third annihilation operator (proceeding from the right) would annihilate it. And if it were ACAA, then normal ordering in $\langle n|H_T|k \rangle$ would convert it to CAAA, with the same effect. The surviving possibilities are summarized in Table 1.

Row 1A of the table contains the pattern AACA of fermion operators. Since we are picking the annihilation operator from the final factor $\psi(\mathbf{x}')$ it must be an electron annihilation operator; and

Case	$ar{\psi}(\mathbf{x}) \ \mathbf{A}(\mathbf{x}) \ \psi(\mathbf{x})$	$\bar{\psi}(\mathbf{x}') \ \mathbf{A}(\mathbf{x}') \ \psi(\mathbf{x}')$
1A	$d_{p's'} a^{\dagger}_{\lambda'} b_{p_1s_1}$	$b^{\dagger}_{p_1s_1}$ a^{\dagger}_{λ} b_{ps}
2A	$d_{p_2s_2} a^{\dagger}_{\lambda} b_{ps}$	$d_{p's'}$ $a^{\dagger}_{\lambda'}$ $d^{\dagger}_{p_2s_2}$
1B	$d_{p's'}$ a^{\dagger}_{λ} $b_{p_1s_1}$	$b^{\dagger}_{p_1s_1}$ $a^{\dagger}_{\lambda'}$ b_{ps}
2B	$d_{p_2s_2} a^{\dagger}_{\lambda'} b_{ps}$	$d_{p's'}$ a^{\dagger}_{λ} $d^{\dagger}_{p_2s_2}$

Table 1. Patterns of creation/annihilation operators that contribute to $e^- - e^+$ annihilation.

since the electron we must annihilate is the incident electron, its mode is (ps), and the operator is b_{ps} , as indicated at the right side of row 1*A*. In the factor $\bar{\psi}(\mathbf{x}')$ we must pick a creation operator, which must be an electron creation operator; we have denoted its mode by (p_1s_1) in the table, and it is an electron in the intermediate state $|k\rangle$. There is also a photon creation operator in $\mathbf{A}(\mathbf{x}')$; it must create one of the final photons, which in row 1*A* we arbitrarily choose to be mode λ . Thus, in row 1A, the intermediate state contains the incident positron, in mode (p's'); the intermediate electron, in mode (p_1s_1) ; and one of the final photons, that in mode λ . The final two annihilation operators in row 1*A* (counting from the right) annihilate the intermediate electron and the incident positron, and a second photon creation operator creates the other outgoing photon.

Row 2A contains the pattern AAAC of fermion operators, in which the first annihilation operator (counting from the right) belongs to $\bar{\psi}(\mathbf{x}')$ and so must be a positron annihilation operator. Therefore it must annihilate the incident positron, and its mode must be (p's') as indicated. Then the creation operator belongs to $\psi(\mathbf{x}')$ and must be a positron creation operator; it must create a positron in the intermediate state, in a mode we denote by (p_2s_2) . The first two fermion operators (counting from the right) are thus $d_{p's'}d^{\dagger}_{p_2s_2}$, as indicated by the table. But when we actually evaluate the matrix element, these must be normal ordered, and replaced with $-d^{\dagger}_{p_2s_2}d_{p's'}$ (with a minus sign). The sign is crucial, because amplitudes add before they are squared to get probabilities, and the wrong sign produces the wrong interference terms. As in row 1A, the final two annihilation operators in AAAC (counting from the right) annihilate the intermediate positron and the incident electron, and a final photon creation operator creates the other final photon.

Finally, rows 1B and 2B differ from 1A and 2A by a swap of the photon labels λ and λ' .

Each row of Table 1 corresponds to a Feynman diagram that contributes to pair annihilation. The diagrams are shown in Fig. 3. In diagrams 1A and 1B the intermediate state contains an electron; in 2A and 2B it contains a positron. The A diagrams and the B diagrams differ only in the order in which the two photons are emitted, that is, by a swap of photon labels λ and λ' .

Instead of wending our way through the combinatorics of creation and annihilation operators, we can directly write down the Feynman diagrams that contribute by combining the diagram fragments that appear in Fig. 2 in such a way as to take us from the given initial state to the given final state. Once we have the diagram it is straightforward to write down the matrix elements, although we



Fig. 3. Feynman diagrams in noncovariant perturbation theory that contribute to the pair annihilation, $e^- + e^+ \rightarrow \gamma \gamma$. must be careful about normal ordering and sign changes.

12. Computing the Matrix Elements

Let us now compute the contribution of diagram 1A to the quantity M, given in Eq. (69). From the diagram we see that the intermediate state contains the initial positron in mode (p's'), the intermediate electron in mode (p_1, s_1) , and the final photon in mode λ . We write it as

$$|k\rangle = a^{\dagger}_{\lambda} b^{\dagger}_{p_1 s_1} d^{\dagger}_{p' s'} |0\rangle, \tag{73}$$

and the sum on k is interpreted as a sum on (p_1, s_1) . We write $p_1 = (E_1, \mathbf{p}_1)$ for the 4-momentum of the intermediate electrons, so that the energy denominator is

$$E_i - E_k = E - E_1 - \omega. \tag{74}$$

The field part of the matrix element $\langle k|H_T|i\rangle$ is obtained from Table 1 or from inspection of the Feynman diagram. It is

$$\langle k|: b_{p_1s_1}^{\dagger} a_{\lambda}^{\dagger} b_{ps}: |i\rangle = \langle 0|d_{p's'} b_{p_1s_1} a_{\lambda} b_{p_1s_1}^{\dagger} a_{\lambda}^{\dagger} b_{ps} b_{ps}^{\dagger} d_{p's'}^{\dagger}|0\rangle,$$
(75)

where in this case the normal ordering does nothing. Because the boson operators a_{λ}^{\dagger} , a_{λ} commute with all fermion operators, they can be migrated adjacent to the vacuum ket on the right, producing $a_{\lambda}a_{\lambda}^{\dagger}|0\rangle = |0\rangle$. In what is left we have

$$\langle 0|d_{p's'}b_{p_1s_1}b^{\dagger}_{p_1s_1}b_{ps}b^{\dagger}_{ps}d^{\dagger}_{p's'}|0\rangle = \langle 0|b_{p_1s_1}b^{\dagger}_{p_1s_1}b_{ps}b^{\dagger}_{ps}d_{p's'}d^{\dagger}_{p's'}|0\rangle,$$
(76)

where we have migrated $d_{p's'}$ four steps to the right, incurring four minus signs. But $d_{p's'}d^{\dagger}_{p's'}|0\rangle = |0\rangle$, so what is left is

$$\langle 0|b_{p_1s_1}b_{p_1s_1}^{\dagger}b_{ps}b_{ps}^{\dagger}|0\rangle = 1,$$
(77)

where we have used $b_{ps}b_{ps}^{\dagger}|0\rangle = |0\rangle$ and $b_{p_1s_1}b_{p_1s_1}^{\dagger}|0\rangle = |0\rangle$, and finally $\langle 0|0\rangle = 1$.





Fig. 4. The matrix element $\langle k|H_T|i\rangle$ conserves 3-momentum at the first vertex (inside the circle), that is, the matrix element vanishes unless $\mathbf{p} = \mathbf{p}_1 + \mathbf{k}$.

Fig. 5. 3-momentum is also conserved at the second vertex, that is, $\langle n|H_T|k\rangle$ vanishes unless $\mathbf{k}' = \mathbf{p}' + \mathbf{p}_1$.

The rest of the matrix element is

$$\langle k|H_T|i\rangle_{1A} = \frac{e}{V}\sqrt{\frac{2\pi}{V}}\int d^3\mathbf{x} \sqrt{\frac{m^2}{E_1E\omega}} \left(\bar{u}_{p_1s_1}\boldsymbol{\gamma}\cdot\hat{\boldsymbol{\epsilon}}^*u_{ps}\right)e^{i\mathbf{x}\cdot(-\mathbf{p}_1-\mathbf{k}+\mathbf{p})}$$

$$= e\sqrt{\frac{2\pi}{V}}\sqrt{\frac{m^2}{E_1E\omega}} \left(\bar{u}_{p_1s_1}\boldsymbol{\gamma}\cdot\hat{\boldsymbol{\epsilon}}^*u_{ps}\right)\delta_{\mathbf{p},\mathbf{p}_1+\mathbf{k}},$$
(78)

where we have written simply $\hat{\epsilon}$ for $\hat{\epsilon}_{\lambda}$. The Kronecker δ shows that momentum is conserved at the first vertex in diagram 1*A*, as shown in Fig. 4.

It is convenient to promote the polarization 3-vector $\hat{\epsilon}$ into a 4-vector by writing

$$\epsilon^{\mu} = (0, \hat{\boldsymbol{\epsilon}}),\tag{79}$$

which is similar to what we did with the electron spin when promoting the spin polarization 3-vector $\hat{\mathbf{s}}$ into the 4-vector s^{μ} . One difference is that the frame in which s^{μ} is purely space-like has an invariant meaning, that is, it is the rest frame of the electron. But the photon has no rest frame, so one can ask what the invariant meaning of Eq. (79) is. The answer is that it has no invariant meaning, rather Coulomb gauge is only valid in one frame, and that is the frame in which ϵ^{μ} is purely space-like. In spite of this the definition is useful, as it allows us to write

$$\not \epsilon = \gamma^{\mu} \epsilon_{\mu} = -\gamma \cdot \hat{\epsilon}, \tag{80}$$

so that

$$\langle k|H_T|i\rangle_{1A} = -e\sqrt{\frac{2\pi}{V}}\sqrt{\frac{m^2}{E_1E\omega}} \left(\bar{u}_{p_1s_1} \not\in^* u_{ps}\right) \delta_{\mathbf{p},\mathbf{p}_1+\mathbf{k}},\tag{81}$$

where $\not\in^*$ means $\gamma^{\mu} \epsilon^*_{\mu}$.

We treat the matrix element $\langle n|H_T|k\rangle$ similarly. The field part of the matrix element is

$$\langle 0|a_{\lambda'}a_{\lambda}: d_{p's'}a_{\lambda'}^{\dagger}b_{p_{1}s_{1}}: a_{\lambda}^{\dagger}b_{p_{1}s_{1}}^{\dagger}d_{p's'}^{\dagger}|0\rangle.$$
(82)

When we normal order and then migrate the boson operators to the right, we obtain $a_{\lambda'}a_{\lambda}a_{\lambda'}^{\dagger}a_{\lambda}^{\dagger}|0\rangle$ on the right, which reduces to simply $|0\rangle$ when we take account of the boson commutation relations and the fact that $\lambda \neq \lambda'$. The latter fact follows from overall energy and momentum conservation, which will be discussed momentarily, which implies $\mathbf{k}' \neq \mathbf{k}$. (This is easily seen in the center of mass frame, in which $\mathbf{k} = -\mathbf{k}'$.) The remaining fermion operators can be reduced as before, showing that the field part of the matrix element (82) is just 1. Evaluating the rest of the matrix element as before, we obtain

The final Kronecker- δ implies 3-momentum conservation at the second vertex of diagram 1A, as shown in Fig. 5.

Putting the pieces from Eqs. (69), (81), (83) and (74) together, we obtain the contribution of diagram 1A to the quantity M,

$$M_{1A} = e^2 \frac{2\pi}{V} \sqrt{\frac{m^2}{EE'\omega\omega'}} \sum_{p_1s_1} \frac{m}{E_1} \frac{(\bar{v}_{p's'} {\not\!\!\!\!\!e'}^{\prime*} u_{p_1s_1})(\bar{u}_{p_1s_1} {\not\!\!\!\!e'}^* u_{ps})}{E - E_1 - \omega} \,\delta_{\mathbf{k}',\mathbf{p}'+\mathbf{p}_1} \,\delta_{\mathbf{p},\mathbf{p}_1+\mathbf{k}}.$$
(84)

The sum on p_1 really means a sum on the 3-momentum \mathbf{p}_1 , which can be done because of the Kronecker deltas. These force

$$\mathbf{p}_1 = \mathbf{p} - \mathbf{k} = \mathbf{k}' - \mathbf{p}',\tag{85}$$

and they leave behind $\delta_{\mathbf{p}+\mathbf{p}',\mathbf{k}+\mathbf{k}'}$ which represents overall conservation of 3-momentum in the Feynman diagram. Although we continue to write \mathbf{p}_1 after the sum is carried out, it stands for the values in Eq. (85). Also, E_1 and p_1 are now understood to mean

$$E_1 = \sqrt{m^2 + |\mathbf{p}_1|^2},\tag{86}$$

and

$$p_1 = (E_1, \mathbf{p}_1).$$
 (87)

As for the spin sum over s_1 , it becomes a positive energy spin projector according to Eqs. (50.38) and (50.146),

$$\sum_{s_1} u_{p_1 s_1} \bar{u}_{p_1 s_1} = \frac{m + \not p_1}{2m} = \Pi_+(p_1).$$
(88)

Altogether, M_{1A} is reduced to

$$M_{1A} = e^2 \frac{2\pi}{V} \,\delta_{\mathbf{p}+\mathbf{p}',\mathbf{k}+\mathbf{k}'} \,\sqrt{\frac{m^2}{EE'\omega\omega'}} \frac{1}{2E_1} \frac{[\bar{v}_{p's'} \not \epsilon''^*(m+\not p_1) \not \epsilon'^* u_{ps}]}{E-E_1-\omega}.$$
(89)

The contribution of diagram 2A to the quantity M can be computed in a similar manner. We summarize the results. The intermediate state is

$$|k\rangle = a^{\dagger}_{\lambda'} d^{\dagger}_{p_2 s_2} b^{\dagger}_{ps} |0\rangle, \tag{90}$$

the energy denominator is

$$E_i - E_k = E' - E_2 - \omega',$$
 (91)

the first matrix element (counting from the right) is

$$\langle k|H_T|i\rangle_{2A} = -e\sqrt{\frac{2\pi}{V}}\sqrt{\frac{m^2}{E'\omega'E_2}} \left(\bar{v}_{p's'} {\not\!\!\!e'}'^* v_{p_2s_2}\right) \delta_{\mathbf{p}',\mathbf{k}'+\mathbf{p}_2},\tag{92}$$

where the minus sign comes from normal ordering and fermion anticommutators, and the second matrix element is

$$\langle n|H_T|k\rangle_{2A} = e\sqrt{\frac{2\pi}{V}}\sqrt{\frac{m^2}{E\omega E_2}} \left(\bar{v}_{p_2s_2} \not\in^* u_{ps}\right) \delta_{\mathbf{k},\mathbf{p}_2+\mathbf{p}}.$$
(93)

Now when we do the p_2 sum we get

$$\mathbf{p}_2 = \mathbf{k} - \mathbf{p} = \mathbf{p}' - \mathbf{k}',\tag{94}$$

and afterwards we understand \mathbf{p}_2 to stand for one of these values. We also understand that

$$E_2 = \sqrt{m^2 + |\mathbf{p}_2|^2},\tag{95}$$

and

$$p_2 = (E_2, \mathbf{p}_2). \tag{96}$$

We also get a Kronecker delta representing overall momentum conservation. As for the s_2 sum, it can be expressed in terms of the negative energy projector, again via Eqs. (50.38) and (50.146),

$$\sum_{s_2} v_{p_2 s_2} \bar{v}_{p_2 s_2} = \frac{\not p_2 - m}{2m} = -\Pi_-(p_2).$$
(97)

Altogether, this gives

$$M_{2A} = -e^2 \frac{2\pi}{V} \,\delta_{\mathbf{p}+\mathbf{p}',\mathbf{k}+\mathbf{k}'} \,\sqrt{\frac{m^2}{EE'\omega\omega'}} \frac{1}{2E_2} \frac{[\bar{v}_{p's'} \,\epsilon''^* (\not\!\!\!p_2 - m) \,\epsilon'^* u_{ps}]}{E' - E_2 - \omega'}.$$
(98)

The amplitudes M_{1B} and M_{2B} are obtained from M_{1A} and M_{2A} , Eqs. (89) and (98), by swapping $\lambda \leftrightarrow \lambda'$, that is, $\mathbf{k} \leftrightarrow \mathbf{k}', \omega \leftrightarrow \omega'$ and $\mathbf{\ell}^* \leftrightarrow \mathbf{\ell}'^*$.

The conservation of 3-momentum at the vertices of the Feynman diagrams means that the 3momentum of the intermediate state is equal to the 3-momentum of both the initial and final states. And the energies of the initial and final states are forced to be equal by the energy-conserving deltafunction in the transition probability (68), in the limit $t \to \infty$. But the energy of the intermediate state is not equal to the energies of the initial and final states. If it were, the energy denominators $E_i - E_k$, which appear in Eqs. (74) and (91), would vanish. The intermediate states arose originally in the Dyson series from a resolution of the identity, which means a sum over all states, regardless of energy.

13. The Feynman Propagator

The amplitudes M_{1A} and M_{2A} can be combined. These amplitudes have many factors in common, so computing the sum boils down to adding the parts that are different, that is, combining the fractions

The algebra of doing this is straightforward. What is interesting is that the result can be written in covariant form.



Fig. 6. Feynman diagrams A and B that appear in covariant perturbation theory. All lines are labelled by their 4-momenta, and 4-momentum is conserved at each vertex. But the 4-momentum of the internal or virtual electron-positron line is not "on mass shell."

We are using noncovariant or "old-fashioned" perturbation theory. It is based on Hamiltonians and the selection of a privileged Lorentz frame. The use of a Hamiltonian, which is the generator of time translations, always implies a privileged Lorentz frame, since the time of one Lorentz frame is not the same as the time of another. Thus, Hamiltonian methods are never covariant (at least this is the usual point of view). Noncovariant perturbation theory is not wrong, but it is not as efficient or as elegant as covariant perturbation theory. In covariant perturbation theory, diagrams 1A and 2A combine into a single diagram A, and likewise diagrams 1B and 2B combine to a single diagram B. These diagrams are illustrated in Fig. 6.

To combine the fractions (99) we first use Eqs. (85) and (94) to express p_2 in terms of p_1 ,

$$p_2 = (E_2, \mathbf{p}_2) = (E_1, -\mathbf{p}_1).$$
 (100)

We then use overall energy conservation, $E + E' = \omega + \omega'$, to express the second denominator in Eq. (99) as

$$E' - E_2 - \omega' = -(E - \omega + E_1).$$
(101)

These allow us to write the expression (99) as

$$\frac{1}{2E_1} \left(\frac{\not p_1 + m}{E - \omega - E_1} + \frac{\not p_2 - m}{E - \omega + E_1} \right) = \frac{1}{2E_1} \frac{(E - \omega)(\not p_1 + \not p_2) + E_1(\not p_1 - \not p_2 + 2m)}{(E - \omega)^2 - E_1^2}.$$
 (102)

But $\not p_1 + \not p_2 = 2E_1\gamma^0$ and $\not p_1 - \not p_2 = -2\mathbf{p}_1 \cdot \boldsymbol{\gamma}$, so the fractions combine to

$$\frac{(E-\omega)\gamma^0 - (\mathbf{p} - \mathbf{k}) \cdot \gamma + m}{(E-\omega)^2 - (\mathbf{p} - \mathbf{k})^2 - m^2} = \frac{\not p - \not k + m}{(p-k)^2 - m^2} = D_F(p-k),$$
(103)

where we define the photon 4-momentum $k = (\omega, \mathbf{k})$, where dot products are understood in the 4-dimensional sense, for example, $(p-k)^2 = (p^{\mu} - k^{\mu})(p_{\mu} - k_{\mu})$, and where for any p,

$$D_F(p) = \frac{\not p + m}{p^2 - m^2}.$$
(104)

The function D_F is called the *Feynman propagator*; it is a kind of Green's function for the free particle Dirac equation.

We can now write out the sum of amplitudes,

$$M_{A} = M_{1A} + M_{2A} = e^{2} \frac{2\pi}{V} \,\delta_{\mathbf{p}+\mathbf{p}',\mathbf{k}+\mathbf{k}'} \,\sqrt{\frac{m^{2}}{EE'\omega\omega'}} \,[\bar{v}_{p's'} \,\epsilon''^{*} D_{F}(p-k) \,\epsilon'' u_{ps}], \tag{105}$$

which is the amplitude associated with diagram A in Fig. 6. We obtain the amplitude for diagram B by swapping photon labels,

$$M_B = e^2 \frac{2\pi}{V} \,\delta_{\mathbf{p}+\mathbf{p}',\mathbf{k}+\mathbf{k}'} \,\sqrt{\frac{m^2}{EE'\omega\omega'}} \,[\bar{v}_{p's'} \,\not\epsilon'^* D_F(p-k') \,\not\epsilon''^* u_{ps}]. \tag{106}$$

To complete the calculation of the cross section we must average over initial electron and positron states (if we are thinking of unpolarized beams). This process is rather tedious so we will stop here.

Courses on quantum field theory normally begin with covariant quantization and perturbation theory, but without any of the background given in these notes which show how those topics are connected with the rest of quantum mechanics or how they developed historically.

Perhaps the biggest question left unanswered by these notes is why the first quantized Dirac equation gives such good results, if we ignore the negative energy solutions. We have explained that the first quantized theory, prior to the introduction of hole theory, is physically incomplete, because of a lack of interpretation of the negative energy solutions. And we have indicated how those interpretational problems are solved by the introduction of field theory. But from a logical standpoint it would be nice to show how and in what sense the first quantized theory is an approximation to field theory. That is an interesting subject that is not usually covered in courses in quantum field theory, which mostly deal with problems that can be solved by the Born approximation. It is perhaps a question that will be covered in a future set of notes.

Problems

1. The purpose of this problem is to derive the equations of motion (50) from the Lagrangian (57), for which the independent fields are ψ , $\bar{\psi}$ and **A**, the last of which is understood to be transverse. The work is carried out in Coulomb gauge.

It is best not to use the Euler-Lagrange equations (51.12), but rather to vary the action $S = \int L dt$ with respect to the independent fields, and to demand that $\delta S = 0$ for all such variations. For practice on this you may see how we obtained Maxwell's equations in Sec. 51.7. Note that when we vary ψ or $\bar{\psi}$, there is a nonvanishing variation $\delta\rho$, because of Eq. (52), which causes a nonvanishing variation $\delta \phi$, because of Eq. (51). If we write S_{Coul} for the Coulomb contribution to the action,

$$S_{\text{Coul}} = \int dt \, \frac{1}{2} \int d^3 \mathbf{x} \, \rho \Phi, \qquad (107)$$

show that under variations of ψ or $\overline{\psi}$ we have

$$\delta S_{\text{Coul}} = \int dt \int d^3 \mathbf{x} \,\delta\rho \,\Phi. \tag{108}$$

Then use this to derive the covariant Dirac equation (50a), driven by the electromagnetic field.

To derive Maxwell's equations driven by the Dirac field, Eq. (50b), we must vary with respect to **A**. Note that if **V** is any vector field and if we demand that

$$\int d^3 \mathbf{x} \, \mathbf{V}(\mathbf{x}) \cdot \delta \mathbf{A}(\mathbf{x}) = 0 \tag{109}$$

for all transverse variations $\delta \mathbf{A}$, then we cannot conclude that $\mathbf{V}(\mathbf{x}) = 0$, only that $\mathbf{V}_{\perp}(\mathbf{x}) = 0$. That is because

$$\int d^3 \mathbf{x} \, \mathbf{V} \cdot \delta \mathbf{A} = \int d^3 \mathbf{x} \left(\mathbf{V}_{\parallel} + \mathbf{V}_{\perp} \right) \cdot \delta \mathbf{A} = \int d^3 \mathbf{x} \, \mathbf{V}_{\perp} \cdot \delta \mathbf{A},\tag{110}$$

where we use Eq. (40.25). Therefore when you vary **A** you will obtain a transverse equation of motion. Show that the longitudinal component of this equation is implied by Coulomb gauge, so that overall the equation of motion is equivalent to Maxwell's equations driven by the Dirac current, Eq. (50b).

This calculation is an example of how Coulomb gauge takes us through noncovariant intermediate steps but produces covariant results. It also shows that the Lagrangian (57) gives the correct equations of motion (for the model we are using).

2. Positrons were first observed by Anderson in 1932 in cosmic ray tracks in a cloud chamber. The experimental apparatus consisted of a series of parallel plates of lead (a high Z material), separated by air gaps. Cosmic rays at the earth's surface are mainly muons, some of which have very high energies. In the Anderson apparatus, a high energy muon, passing close to a nucleus of lead and being accelerated in the electric field, produced a gamma ray. This process is called *Bremsstrahlung* (German for "braking radiation"). The gamma ray, in turn, passing near a lead nucleus in one of the lower plates, produced an electron-positron pair.

To model this process, we add a term to the perturbing Hamiltonian $H_1 = H_{\text{Coul}} + H_T$, given by Eqs. (64), to account for the interaction with the nucleus,

$$H_{\text{ext}} = \int d^3 \mathbf{x} : \rho(\mathbf{x}) \Phi_{\text{ext}}(\mathbf{x}) := -e \int d^3 \mathbf{x} : \bar{\psi}(\mathbf{x}) \gamma^0 \Phi_{\text{ext}}(\mathbf{x}) \psi(\mathbf{x}) :.$$
(111)

This is exactly how we treated the field of the nucleus Mott scattering (see Eq. (12) and Sec. 3). In effect, we are dividing the electromagnetic field into a *c*-number field (from the nucleus) plus the quantized field. If we wish to model the lead nucleus as a point charge, we can take $\Phi_{\text{ext}}(\mathbf{x}) = Ze/r$, but for this problem we will leave Φ_{ext} unspecified. Thus, the perturbing Hamiltonian is now

$$H_1 = H_{\text{Coul}} + H_T + H_{\text{ext}}.$$
(112)

(a) Consider a process in which a photon passes close to a nucleus, and an electron-positron pair is produced. (This process cannot happen in free space because of energy-momentum conservation.) Let (ps) be the 4-momentum and spin of the outgoing electron, let (p's') be the 4-momentum and spin of the outgoing positron, let $\lambda = (\mathbf{k}, \boldsymbol{\epsilon})$ be the mode of the incident photon. Also write

$$p^{\mu} = (E, \mathbf{p}),$$

 $p'^{\mu} = (E', \mathbf{p}'),$
 $k^{\mu} = (\omega, \mathbf{k}).$ (113)

Find and draw all Feynman diagrams which contribute to this process at lowest order in $\alpha = e^2$. Count the external potential Φ as containing one power of e (since we are thinking of $\Phi = Ze/r$). Indicate the interaction with the nucleus by an \times drawn next to an electron or positron line, as in Fig. 1.

(b) Write the transition probability as in Eq. (68), which defines M. Pick out two Feynman diagrams which differ from one another only in the time ordering of the creation of the outgoing electron and positron. Work out in detail the contribution to M from each of these Feynman diagrams. Express your answer in terms of the Fourier transform of the potential $\tilde{\Phi}(\mathbf{q})$, as we did with Mott scattering.

(c) Combine the two terms, and express the result in terms of the Feynman electron propagator (104).

The steps remaining to convert M into a cross section are straightforward but tedious so we will stop at this point.