## Physics 209

Fall 2002
Homework 9

## Due Saturday, November 1 at 5:00pm.

Note: Due to the exam in 221A, this homework is due on Saturday, November 1 instead of Friday, October 31. However, the building is only open during limited hours on Saturday. If you do not have a key, you will have to make arrangements to have your homework delivered on time, or else turn it in on Friday.

1. In class we derived the $3+1$ version of the relativistic Lagrangian for a particle in a given electric and magnetic field,

$$
\begin{equation*}
L(\mathbf{x}, \dot{\mathbf{x}})=-m c^{2} \sqrt{1-|\dot{\mathbf{x}}|^{2} / c^{2}}-e \Phi+\frac{e}{c} \dot{\mathbf{x}} \cdot \mathbf{A} . \tag{1}
\end{equation*}
$$

(a) Show that the Euler-Lagrange equations give the correct equations of motion.
(b) Show that this Lagrangian is regular (you can solve for $\dot{\mathbf{x}}$ in terms of $\mathbf{p}$ ), and find the Hamiltonian $H(\mathbf{x}, \mathbf{p})$
(c) Show that Hamilton's equations give the correct equations of motion.
2. Noether's theorem for Lagrangians $L(q, \dot{q})$ was described in class as follows. We let $q \rightarrow Q(q, \epsilon)$ be a mapping of configuration space onto itself. We say that this mapping is a symmetry of the Lagrangian if

$$
\begin{equation*}
L\left(Q(q, \epsilon), \frac{d Q(q, \epsilon)}{d t}\right)=L(q, \dot{q})+\frac{d \Lambda(q, \epsilon)}{d t} \tag{2}
\end{equation*}
$$

for some function $\Lambda$, where $q, Q$ etc are abbreviations for multicomponent coordinate vectors, $q_{i}, Q_{i}$, etc. Also, in Eq. (2) $d Q / d t$ and $d \Lambda / d t$ are abbreviations for

$$
\begin{equation*}
\frac{d Q_{i}(q, \epsilon)}{d t}=\sum_{j} \frac{\partial Q_{i}}{\partial q_{j}} \dot{q}_{j} \tag{3}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{d \Lambda(q, \epsilon)}{d t}=\sum_{i} \frac{\partial \Lambda}{\partial q_{i}} \dot{q}_{i}, \tag{4}
\end{equation*}
$$

so that the Lagrangian before and after the transformation is a function of $q$ and $\dot{q}$. Equation (2) means that the equations of motion are the same before and after the symmetry transformation.

As explained in class, we only use this symmetry property for infinitesimal $\epsilon$, which means we expand out to first order in $\epsilon$ to obtain an identity involving the Lagrangian, or else we differentiate Eq. (2) with respect to $\epsilon$ and set $\epsilon=0$, which is the same thing. This gives the identity,

$$
\begin{equation*}
\sum_{i}\left(\frac{\partial L}{\partial q_{i}} F_{i}+\frac{\partial L}{\partial \dot{q}_{i}} \frac{d F_{i}}{d t}\right)=\frac{d G}{d t} \tag{5}
\end{equation*}
$$

where

$$
\begin{equation*}
F_{i}(q)=\left.\frac{\partial Q_{i}(q, \epsilon)}{\partial \epsilon}\right|_{\epsilon=0} \tag{6}
\end{equation*}
$$

and

$$
\begin{equation*}
G(q)=\left.\frac{\partial \Lambda(q, \epsilon)}{\partial \epsilon}\right|_{\epsilon=0} \tag{7}
\end{equation*}
$$

Equation (5) is just an identity satisfied by the Lagrangian, due to the symmetry. But now we use the Euler-Lagrange equations,

$$
\begin{equation*}
\frac{d}{d t}\left(\frac{\partial L}{\partial \dot{q}_{i}}\right)=\frac{\partial L}{\partial q_{i}} \tag{8}
\end{equation*}
$$

to write Eq. (5) in the form,

$$
\begin{equation*}
\frac{d}{d t}\left(\sum_{i} F_{i} p_{i}-G\right)=0 \tag{9}
\end{equation*}
$$

This gives us the conserved quantity corresponding to the symmetry.
(a) Consider the relativistic motion of a charged particle of charge $e$ in a constant, uniform magnetic field, $\mathbf{B}=B_{0} \hat{\mathbf{z}}$. Use the covariant Lagrangian,

$$
\begin{equation*}
L\left(x^{\mu}, \frac{d x^{\mu}}{d \sigma}\right)=m c \sqrt{\frac{d x^{\mu}}{d \sigma} \frac{d x_{\mu}}{d \sigma}}-\frac{e B_{0}}{c} x \frac{d y}{d \sigma} \tag{10}
\end{equation*}
$$

where as explained in class $t$ is regarded as a 4 th $q$ and $\sigma$ is an arbitrary parameter of the orbits. Equation (10) effectively uses the vector potential, $\mathbf{A}=B_{0} x \hat{\mathbf{y}}$. Find the constants of motion associated with translations in the $x, y, z$ and $t$ directions. Find the constant of motion associated with rotations in the $x-y$ plane. Express your constants of motion in terms of the kinetic (not canonical) momentum, or equivalently in terms of the world velocity. Show that the constants associated with $x$ and $y$ displacements are closely related to the guiding center position.

Notational suggestion: Since the covariant, kinetic 4-momentum $p_{\mu}$ has components,

$$
\begin{equation*}
p_{\mu}=\binom{E / c}{-\mathbf{p}} \tag{11}
\end{equation*}
$$

use the notation $p_{1}, p_{2}, p_{3}$ for $p_{\mu}$ with $\mu=1,2,3$, and $p_{x}, p_{y}, p_{z}$ for the usual components of $\mathbf{p}$, so that $p_{x}=-p_{1}$, etc.
(b) This is probably the hardest part of this homework, so you might want to do it last. Consider the motion of a nonrelativistic charged particle of charge $e$ in the field of a magnetic monopole,

$$
\begin{equation*}
\mathbf{B}(\mathbf{x})=g \frac{\mathbf{x}}{r^{3}} \tag{12}
\end{equation*}
$$

where $g$ is the strength of the monopole and $r=|\mathbf{x}|$. Assume the monopole is infinitely massive and located at the origin. The nonrelativistic Lagrangian for this system is

$$
\begin{equation*}
L=\frac{m}{2}|\dot{\mathbf{x}}|^{2}+\frac{e}{c} \dot{\mathbf{x}} \cdot \mathbf{A}(\mathbf{x}) \tag{13}
\end{equation*}
$$

where $\mathbf{B}=\nabla \times \mathbf{A}$. It is possible to work out explicit formulas for $\mathbf{A}$ which produce the magnetic field (12), but the formulas are ugly, so we will not attempt to do that. Instead, we will just use the property $\mathbf{B}=\nabla \times \mathbf{A}$, with $\mathbf{B}$ given by Eq. (12).

Consider rotations about the axis $\hat{\mathbf{n}}$, so that for small angles of rotation $\theta$, $\mathbf{x}$ is replaced by $\mathbf{x}+\theta(\hat{\mathbf{n}} \times \mathbf{x})$. Show that such rotations are a symmetries of the Lagrangian (13), and find the corresponding vector of conserved quantities (a vector since $\hat{\mathbf{n}}$ is arbitrary). Hint: the conserved quantities will be gauge-invariant, that is, you should be able to find an expression for them that does not involve $\mathbf{A}$.
3. Fermat's principle in optics says that light rays $\mathbf{x}(\lambda)$ are stationary paths of the following functional, the "optical path,"

$$
\begin{equation*}
P[\mathbf{x}(\lambda)]=c \int d t=\int n(\mathbf{x}) d s=\int_{\lambda_{0}}^{\lambda_{1}} n(\mathbf{x})\left|\frac{d \mathbf{x}}{d \lambda}\right| d \lambda \tag{14}
\end{equation*}
$$

where $n(\mathbf{x})$ is the index of refraction, $\mathbf{x}=(x, y, z)$, and $\lambda$ is an arbitrary parameter.
Suppose the medium is one in which $n(\mathbf{x})$ only depends on $z$, so that the system has translational symmetry in the $x$ - and $y$-directions. For example, we might have a discontinuous change from one medium to another, with the interface being at $z=$ const. Show that Snell's law is a consequence of Noether's theorem.
4. In class we modelled the transverse vibrations of a string as a discrete set of $N$ masses $m$ spaced at intervals of $\Delta x$ with Lagrangian,

$$
\begin{equation*}
L=\sum_{i=1}^{N} \frac{m}{2} \dot{y}_{i}^{2}-\sum_{i=1}^{N+1} \frac{k}{2}\left(y_{i}-y_{i-1}\right)^{2}, \tag{15}
\end{equation*}
$$

where $k$ is the spring constant for the spring connecting mass $i$ and mass $i-1$. All springs have the same $k$. Then we took the limit in which $N \rightarrow \infty, m=\rho \Delta x, k=\kappa / \Delta x$, where $\rho$ is the mass per unit length of the string and $\kappa$ is the spring constant of the whole string times the length of the string. This gave the field Lagrangian,

$$
\begin{equation*}
L=\int d x \mathcal{L} \tag{16}
\end{equation*}
$$

where the Lagrangian density is

$$
\begin{equation*}
\mathcal{L}=\frac{\rho}{2}\left(\frac{\partial y}{\partial t}\right)^{2}-\frac{\kappa}{2}\left(\frac{\partial y}{\partial x}\right)^{2} . \tag{17}
\end{equation*}
$$

Construct the classical Hamiltonian $H$ for the discretized version of the vibrating string. Take the limit $N \rightarrow \infty$ and show that $H$ can be written as the $x$ integral of a Hamiltonian density $\mathcal{H}$, and find an expression for $\mathcal{H}$. The Hamiltonian density $\mathcal{H}$ is interpreted as the energy density of the system.

Show that there exists a function $S$, depending on $y(x, t)$ and/or its various derivatives, such that

$$
\begin{equation*}
\frac{\partial \mathcal{H}}{\partial t}+\frac{\partial S}{\partial x}=0 . \tag{18}
\end{equation*}
$$

This is a one-dimensional continuity equation, and $S$ is interpreted as the energy flux of the system (like the Poynting vector in electromagnetism).

